

Stochastic particle production in the early universe

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Based on 1705.????, with

M. Amin, O. Wen (Rice U.)

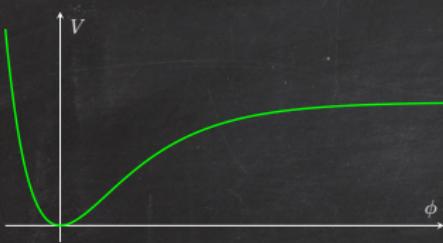
H. Xie (U. Wisconsin)



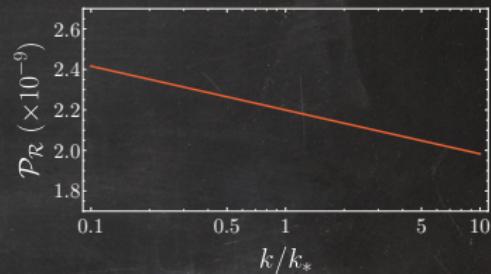
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Motivation

Cosmological inflation is the early period of accelerated expansion, $a \sim e^{Ht}$

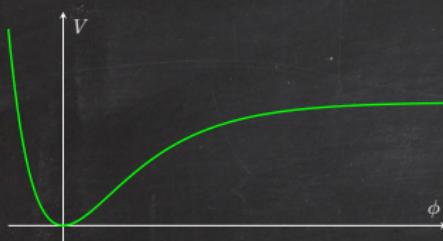


$$\frac{\phi \rightarrow \phi + \delta\phi}{g \rightarrow g + \delta g} \rightarrow$$

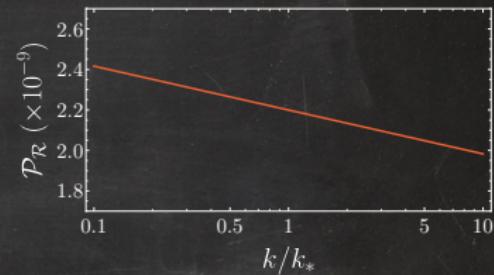


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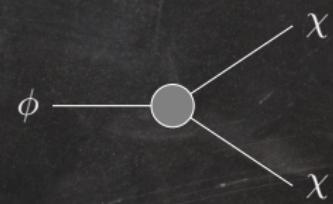
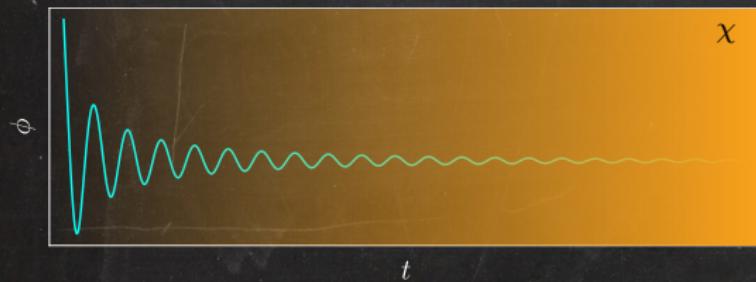
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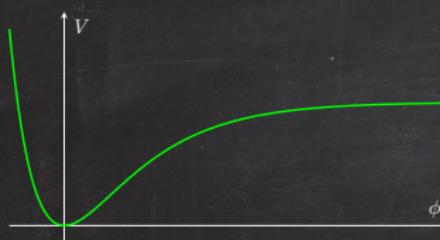


When inflation ends, the Universe reheats...

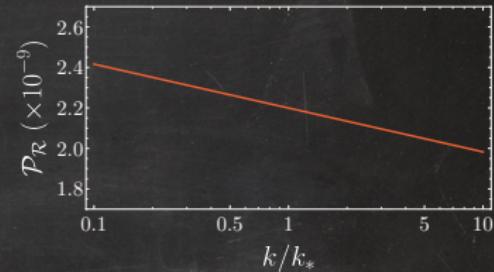


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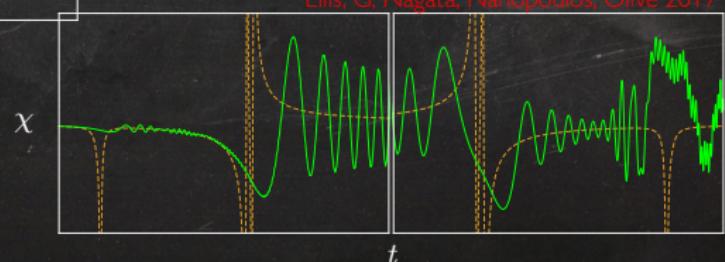
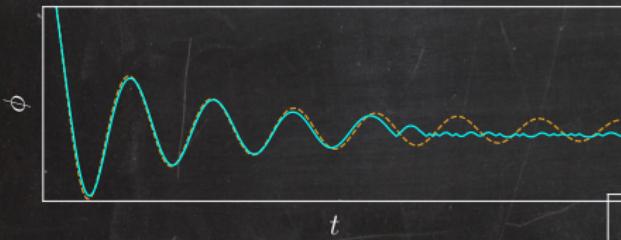
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When inflation ends, the Universe reheats...in a potentially complicated way



Ellis, G, Nagata, Nanopoulos, Olive 2017

Stochastic Particle Production

Consider N_f coupled (scalar) fields. Assume the evolution of fluctuations contains localized non-adiabatic events with random strengths at random intervals, and that the fields are otherwise free

$$\left[\mathbb{I} \partial_\tau^2 + \omega^2 + \mathbf{m}^s(\tau) \right] \cdot \chi(\tau, \mathbf{k}) = 0, \quad \omega_a^2 = k^2 + m_a^2$$



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After the j -th event,

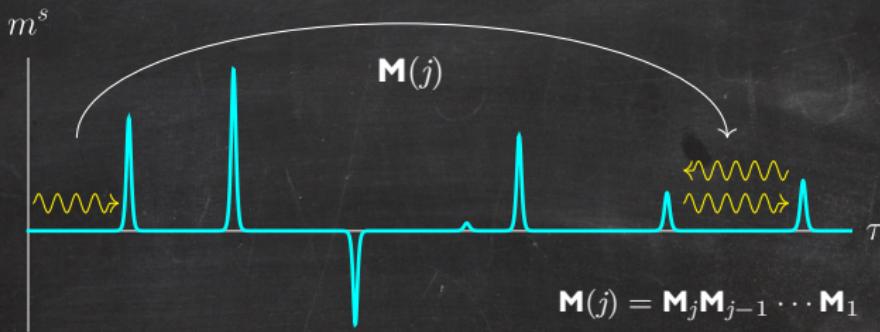
$$\chi_j^a(\tau) \equiv \frac{1}{\sqrt{2\omega_a}} \left[\beta_j^a e^{i\omega_a \tau} + \alpha_j^a e^{-i\omega_a \tau} \right],$$

$$\begin{pmatrix} \beta_j \\ \alpha_j \end{pmatrix} = \mathbf{M}_j \begin{pmatrix} \beta_{j-1} \\ \alpha_{j-1} \end{pmatrix}$$

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A random walk (with drift) for the occupation number

$$n_a(j) = \frac{1}{2\omega_a} \left(|\dot{\chi}_j^a|^2 + \omega_a^2 |\chi_j^a|^2 \right) - \frac{1}{2} = |\beta_j^a|^2$$

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Randomness and non-adiabaticity are encoded in \mathbf{M} \Rightarrow define $P(\mathbf{M}; \tau)$

$$\partial_\tau P(\mathbf{M}; \tau) = -\partial_{\mathbf{M}} \left[\frac{\langle \delta \mathbf{M} \rangle_{\mathbf{M}_2}}{\delta \tau} P(\mathbf{M}; \tau) \right] + \frac{1}{2!} \partial_{\mathbf{M}}^2 \left[\frac{\langle \delta \mathbf{M}^2 \rangle_{\mathbf{M}_2}}{\delta \tau} P(\mathbf{M}; \tau) \right]$$

A general transfer matrix can be parametrized as

$$\mathbf{M} = \begin{pmatrix} \mathbf{u} & 0 \\ 0 & \mathbf{u}^* \end{pmatrix} \begin{pmatrix} \sqrt{1+\mathbf{n}} & \sqrt{\mathbf{n}} \\ \sqrt{\mathbf{n}} & \sqrt{1+\mathbf{n}} \end{pmatrix} \begin{pmatrix} \mathbf{v} & 0 \\ 0 & \mathbf{v}^* \end{pmatrix}$$

where $\mathbf{u}, \mathbf{v} \in U(N_f)$, and $\mathbf{n} = \text{diag}(n_1, n_2, \dots) \Rightarrow N_f(2N_f + 1)$ variables in FP equation!

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The Maximum Entropy Ansatz

Assume the building block P maximizes the Shannon entropy

$$S[P] = - \int P(\mathbf{M}; \delta\tau) \ln P(\mathbf{M}; \delta\tau) d\mathbf{M}$$

subject to the constraints:

- The local mean particle production rate is known and fixed, $\mu_j \equiv \frac{1}{N_f} \frac{\langle n_j \rangle_{\delta\tau}}{\delta\tau}$
- Coarse-grained continuity, $\mathbf{M}_{\tau+\delta\tau} \xrightarrow{\delta\tau \rightarrow 0} \mathbf{M}_\tau$

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The Maximum Entropy Ansatz

Then (Mello, Pereyra, Kumar 1988; Amin, Baumann 2016),

- P is flat (Haar) wrt \mathbf{u} , $dP(\{\mathbf{u}, \mathbf{n}, \mathbf{v}\}) = P(\{\mathbf{n}, \mathbf{v}\}) d\mu(\mathbf{u})$
- A closed set of equations for the moments of $n = \sum_a n_a$ is obtained,

$$\partial_\tau \langle \ln(1+n) \rangle \xrightarrow{\tau \rightarrow \infty} \frac{2N_f}{N_f + 1} \mu$$

i.e. exponential growth

Exact Results

Consider the approximation

$$m_{ab}^s(\tau) = 2\sqrt{\omega_a \omega_b} \sum_{j=1}^{N_s} \Lambda_{ab}(\tau_j) \delta(\tau - \tau_j),$$

where τ_j are uniformly distributed, and

$$\langle \Lambda_{ab} \rangle = 0, \quad \langle \Lambda_{ab} \Lambda_{cd} \rangle = \sigma_{ab}^2 (\delta_{ac} \delta_{bd} + \delta_{ad} \delta_{bc})$$

The transfer matrix takes the form

$$\mathbf{M}_j = \mathbb{1} + i \underbrace{\begin{pmatrix} \mathbf{a}_j^* & 0 \\ 0 & \mathbf{a}_j \end{pmatrix} \begin{pmatrix} \boldsymbol{\Lambda}_j & \boldsymbol{\Lambda}_j \\ -\boldsymbol{\Lambda}_j & -\boldsymbol{\Lambda}_j \end{pmatrix} \begin{pmatrix} \mathbf{a}_j & 0 \\ 0 & \mathbf{a}_j^* \end{pmatrix}}_{\mathbf{m}_j}, \quad \mathbf{a}_j \equiv \text{diag}(e^{i\omega_1 \tau_j}, e^{i\omega_2 \tau_j}, \dots)$$

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Will focus on the total occupation number. Define $\mathbf{R} = \mathbf{M}\mathbf{M}^\dagger$:

$$n(j) = \frac{1}{4} \text{Tr} [\mathbf{M}(j) \mathbf{M}^\dagger(j) - \mathbb{1}] \equiv \frac{1}{4} \text{Tr} [\mathbf{R}(j) - \mathbb{1}]$$

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Single field

Only two parameters, n and $u = e^{i\phi}$. Computation is straightforward,

$$\langle \delta n \rangle = (2n+1) \sigma^2$$

$$\langle \delta \phi \rangle = 0$$

$$\langle \delta n \delta n \rangle = 2n(n+1)\sigma^2$$

$$\langle \delta n \delta \phi \rangle = 0$$

$$\langle \delta \phi \delta \phi \rangle = \frac{\sigma^2}{8n(n+1)}(12n^2 + 12n + 1)$$

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The FP equation is

$$\frac{1}{\sigma^2} \frac{\partial}{\partial \tau} P(n; \tau) = \frac{\partial}{\partial n} \left[n(n+1) \frac{\partial}{\partial n} P(n; \tau) \right]$$

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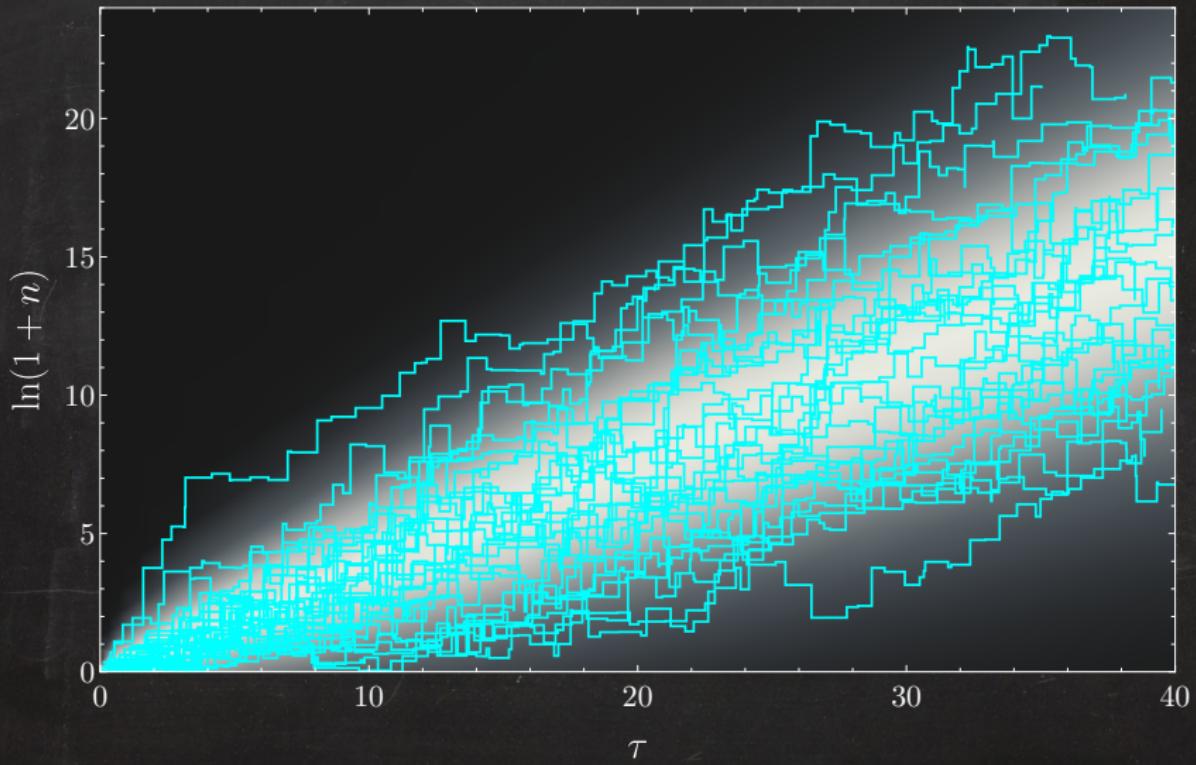
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with solution ($n \gg 1$)

$$P(n; \tau) dn = \frac{1}{\sqrt{4\pi\sigma^2\tau}} \exp \left[-\frac{(\ln n - \sigma^2\tau)^2}{4\sigma^2\tau} \right] d\ln n$$

$$\Rightarrow n = e^{\sigma^2\tau} - 1$$



Two fields

Six parameters now, n_1 , n_2 and

$$\mathbf{u}(\phi, \theta, \psi, \varphi) = e^{-\frac{i}{2}\phi} \begin{bmatrix} \cos \frac{\theta}{2} e^{-\frac{i}{2}(\varphi+\psi)} & -\sin \frac{\theta}{2} e^{-\frac{i}{2}(\varphi-\psi)} \\ \sin \frac{\theta}{2} e^{\frac{i}{2}(\varphi-\psi)} & \cos \frac{\theta}{2} e^{\frac{i}{2}(\varphi+\psi)} \end{bmatrix}$$

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⇒ need 27 correlators

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Let

$$\langle (\Lambda_{11})^2 \rangle_{\delta\tau} = \sigma_1^2, \quad \langle (\Lambda_{22})^2 \rangle_{\delta\tau} = \sigma_2^2, \quad \langle (\Lambda_{12})^2 \rangle_{\delta\tau} = \langle (\Lambda_{21})^2 \rangle_{\delta\tau} = \sigma_{\perp}^2.$$

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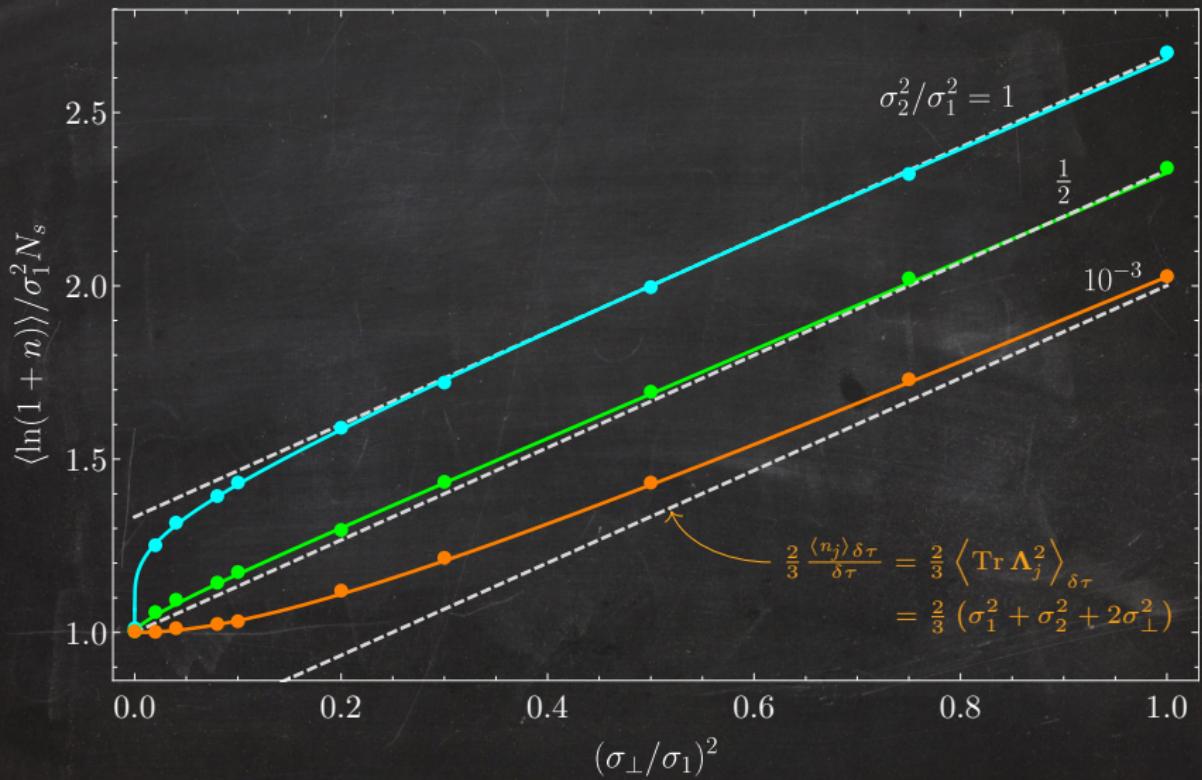
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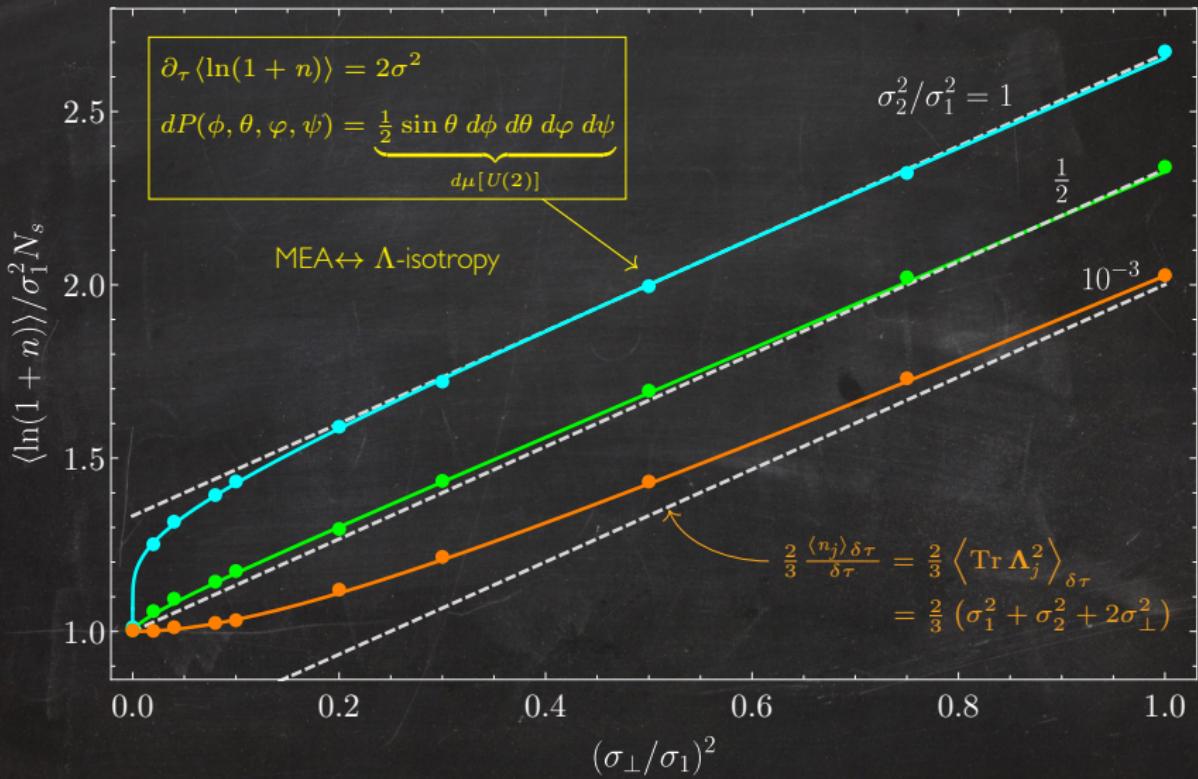
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$$\begin{aligned} \langle \delta f_1^{(1)} \delta f_1^{(1)} \rangle &= \tilde{f}_1^2 \gamma_1(\theta), \\ \langle \delta f_1^{(1)} \delta f_2^{(1)} \rangle &= 2\tilde{f}_1 \tilde{f}_2 \cos(2\psi) \gamma_3(\theta), \\ \langle \delta f_1^{(1)} \delta \theta^{(1)} \rangle &= -\frac{2\tilde{f}_1}{\Delta f} [\tilde{f}_1 + \tilde{f}_2 \cos(2\psi)] \gamma_4(\theta), \\ \langle \delta f_1^{(1)} \delta \psi^{(1)} \rangle &= -\tilde{f}_1 \sin(2\psi) \left(\frac{\tilde{f}_2}{\tilde{f}_1} \gamma_2(\theta) - 2\frac{\tilde{f}_2}{\Delta f} \gamma_4(\theta) \cot\theta \right), \\ \langle \delta f_1^{(1)} \delta \varphi^{(1)} \rangle &= -2\frac{\tilde{f}_1 \tilde{f}_2}{\Delta f} \sin(2\psi) \gamma_4(\theta) \csc\theta, \\ \langle \delta f_1^{(1)} \delta \phi^{(1)} \rangle &= \tilde{f}_1 \frac{\tilde{f}_2}{\tilde{f}_1} \sin(2\psi) \gamma_3(\theta), \\ \langle \delta f_2^{(1)} \delta f_2^{(1)} \rangle &= \tilde{f}_2^2 \gamma_2(\theta), \\ \langle \delta f_2^{(1)} \delta \theta^{(1)} \rangle &= -\frac{2\tilde{f}_2}{\Delta f} [\tilde{f}_2 + \tilde{f}_1 \cos(2\psi)] \gamma_5(\theta), \\ \langle \delta f_2^{(1)} \delta \psi^{(1)} \rangle &= -\tilde{f}_2 \sin(2\psi) \left(\frac{\tilde{f}_1}{\tilde{f}_2} \gamma_2(\theta) - 2\frac{\tilde{f}_1}{\Delta f} \gamma_4(\theta) \cot\theta \right), \\ \langle \delta f_2^{(1)} \delta \varphi^{(1)} \rangle &= -2\frac{\tilde{f}_1 \tilde{f}_2}{\Delta f} \sin(2\psi) \gamma_5(\theta) \csc\theta, \\ \langle \delta f_2^{(1)} \delta \phi^{(1)} \rangle &= -\tilde{f}_2 \frac{\tilde{f}_1}{\tilde{f}_2} \sin(2\psi) \gamma_3(\theta), \\ \langle \delta \theta^{(1)} \delta \theta^{(1)} \rangle &= 2\sigma_1^2 + \frac{1}{\Delta f^2} (\tilde{f}_1 + \tilde{f}_2 + 2\tilde{f}_1 \tilde{f}_2 \cos(2\psi)), \\ \langle \delta \theta^{(1)} \delta \psi^{(1)} \rangle &= \frac{1}{\Delta f} \left[\frac{\tilde{f}_1 \tilde{f}_2}{\tilde{f}_1 + \tilde{f}_2} \gamma_4(\theta) + \frac{\tilde{f}_1 \tilde{f}_2}{\tilde{f}_2} \gamma_5(\theta) \right] \sin(2\psi) - \cos\theta \langle \delta \theta^{(1)} \delta \varphi^{(1)} \rangle, \\ \langle \delta \theta^{(1)} \delta \varphi^{(1)} \rangle &= \frac{2\tilde{f}_1 \tilde{f}_2}{\Delta f^2} \sin(2\psi) \gamma_6(\theta) \csc\theta, \\ \langle \delta \theta^{(1)} \delta \phi^{(1)} \rangle &= \frac{1}{\Delta f} \left[\tilde{f}_1^2 \gamma_1(\theta) - (\tilde{f}_1 \tilde{f}_2 - 1) \gamma_3(\theta) \right], \end{aligned}$$

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$$\begin{aligned} \langle \delta \theta^{(1)} \delta \phi^{(1)} \rangle &= \frac{1}{4\Delta f} \left[\left(\frac{\tilde{f}_1}{\tilde{f}_1} \frac{\tilde{f}_2}{\tilde{f}_1} - \frac{\tilde{f}_2}{\tilde{f}_2} \frac{\tilde{f}_1}{\tilde{f}_2} \right) (\sigma_1^2 - \sigma_2^2) \sin\theta + 4 \left(\frac{\tilde{f}_1}{\tilde{f}_1} \frac{\tilde{f}_2}{\tilde{f}_1} + \frac{\tilde{f}_2}{\tilde{f}_2} \frac{\tilde{f}_1}{\tilde{f}_2} \right) \right] \sin(2\psi) \\ \langle \delta \psi^{(1)} \delta \psi^{(1)} \rangle &= \frac{1}{4} \left[\left(\frac{\tilde{f}_1}{\tilde{f}_1} \right)^2 \gamma_1(\theta) + \left(\frac{\tilde{f}_2}{\tilde{f}_2} \right)^2 \gamma_2(\theta) \right] - \frac{\tilde{f}_1 \tilde{f}_2}{\Delta f} (\sigma_1^2 + \sigma_2^2) + 2\sin^2\theta \sigma_{\perp}^2 - \cos\theta \left(2 \langle \delta \psi^{(1)} \delta \varphi^{(1)} \rangle + \cos\theta \langle \delta \varphi^{(1)} \delta \varphi^{(1)} \rangle \right), \\ \langle \delta \psi^{(1)} \delta \varphi^{(1)} \rangle &= -\frac{1}{\Delta f} \left[\frac{\tilde{f}_1 \tilde{f}_2}{\tilde{f}_1 + \tilde{f}_2} \gamma_4(\theta) + \frac{\tilde{f}_1 \tilde{f}_2}{\tilde{f}_2} \gamma_5(\theta) \right] \sin(2\psi) - \frac{\tilde{f}_1 \tilde{f}_2}{\Delta f} (\sigma_1^2 + \sigma_2^2) - 3\sigma_{\perp}^2 \left[\cos\theta - \cos\theta \langle \delta \varphi^{(1)} \delta \varphi^{(1)} \rangle \right. \\ &\quad \left. - \langle \delta \varphi^{(1)} \delta \phi^{(1)} \rangle \cos\theta \right], \\ \langle \delta \psi^{(1)} \delta \phi^{(1)} \rangle &= \frac{1}{4} \left[\left(\frac{\tilde{f}_1}{\tilde{f}_1} - \frac{\tilde{f}_2}{\tilde{f}_2} \right) \cos\theta - \frac{1}{2} \left[\frac{\tilde{f}_1}{\tilde{f}_1} + \frac{\tilde{f}_2}{\tilde{f}_2} - 2\frac{\tilde{f}_1 \tilde{f}_2}{\tilde{f}_1 + \tilde{f}_2} \cos(2\psi) \right] \gamma_3(\theta) \right], \\ \langle \delta \varphi^{(1)} \delta \varphi^{(1)} \rangle &= \frac{1}{2} \left[\left(\frac{\tilde{f}_1}{\tilde{f}_1} + \frac{\tilde{f}_2}{\tilde{f}_2} \right) (\sigma_1^2 + \sigma_2^2 - \sigma_{\perp}^2) \sin\theta - \frac{1}{2} \sin(2\theta) (\sigma_1^2 + \sigma_2^2 - 3\sigma_{\perp}^2) \right. \\ &\quad \left. + \frac{1}{\Delta f} \left(\langle \delta f_2^{(1)} \delta \theta^{(1)} \rangle - \langle \delta f_1^{(1)} \delta \theta^{(1)} \rangle \right) + \frac{1}{4} \sin(2\theta) \langle \delta \psi^{(1)} \delta \varphi^{(1)} \rangle \right. \\ &\quad \left. - \frac{1}{8} \left(\frac{\tilde{f}_1}{\tilde{f}_1} + \frac{\tilde{f}_2}{\tilde{f}_2} \right) (\sigma_1^2 + \sigma_2^2 - 2\sigma_{\perp}^2) \sin^2\theta \sin(2\psi) - \frac{1}{2} \left(\frac{\tilde{f}_1}{\tilde{f}_1} \langle \delta f_1^{(1)} \delta \phi^{(1)} \rangle - \frac{\tilde{f}_2}{\tilde{f}_2} \langle \delta f_2^{(1)} \delta \phi^{(1)} \rangle \right) \right. \\ &\quad \left. - \frac{1}{2} \left(\frac{\tilde{f}_1}{\tilde{f}_1} \langle \delta f_1^{(1)} \delta \varphi^{(1)} \rangle + \frac{\tilde{f}_2}{\tilde{f}_2} \langle \delta f_2^{(1)} \delta \varphi^{(1)} \rangle \right) \cos\theta + \frac{1}{2} \sin\theta \langle \delta \theta^{(1)} \delta \varphi^{(1)} \rangle - \cos\theta \langle \delta \varphi^{(1)} \delta \varphi^{(1)} \rangle \right. \\ &\quad \left. - \frac{1}{2} \left(\frac{\tilde{f}_1}{\tilde{f}_1} - \frac{\tilde{f}_2}{\tilde{f}_2} \right) \left[\frac{1}{4} \left(\langle \delta \varphi^{(1)} \delta \varphi^{(1)} \rangle \sin^2\theta - \langle \delta \theta^{(1)} \delta \theta^{(1)} \rangle \right) \sin 2\psi + \frac{1}{2} \langle \delta \varphi^{(1)} \delta \theta^{(1)} \rangle \sin\theta \cos 2\psi \right] \right], \\ \langle \delta \varphi^{(1)} \delta \theta^{(1)} \rangle &= \frac{1}{\Delta f} \left[\langle \delta f_2^{(1)} \delta \varphi^{(1)} \rangle - \langle \delta f_1^{(1)} \delta \varphi^{(1)} \rangle \right] - \cot\theta \langle \delta \theta^{(1)} \delta \varphi^{(1)} \rangle, \\ \langle \delta \varphi^{(2)} \rangle &= 0 \end{aligned}$$





N_f fields

$N_f(N_f + 1)$ parameters now, n_1, n_2, \dots, n_{N_f} and (Tilma, Sudarshan 2002)

$$\mathbf{u} = \left(\prod_{2 \leq k \leq N} \mathbf{A}(k) \right) \cdot [SU(N-1)] \cdot e^{i\lambda_{N^2-1}\alpha_{N^2-1}}, \quad \mathbf{A}(k) = e^{i\lambda_3\alpha_{(2k-3)}} e^{i\lambda_{(k-1)^2+1}\alpha_{2(k-1)}}$$

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\Rightarrow need $\mathcal{O}(N_f^4)$ correlators!

N_f fields

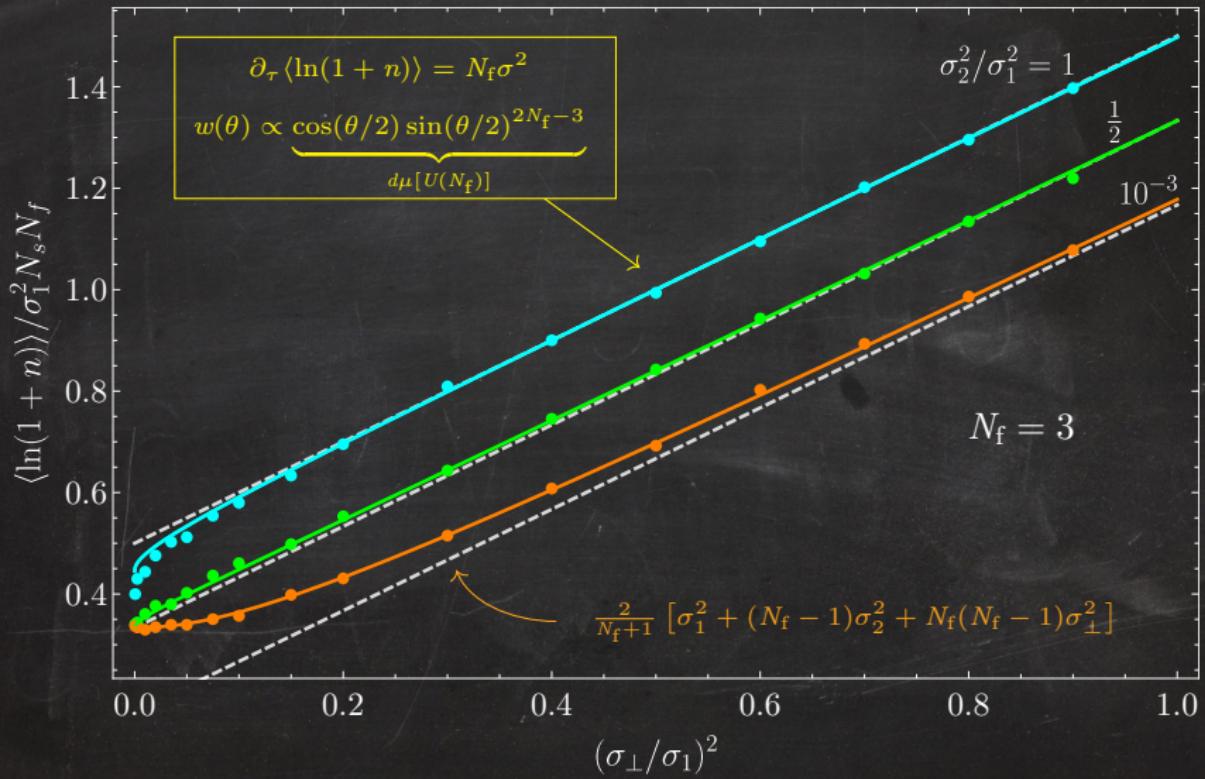
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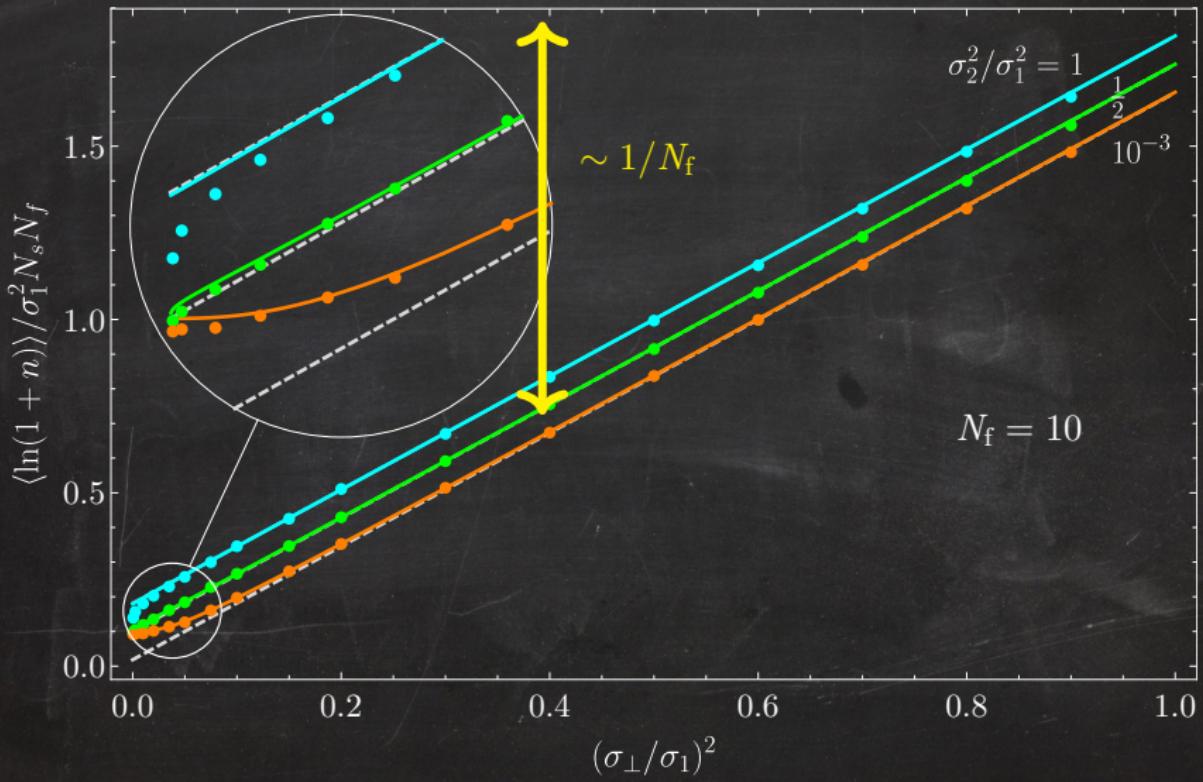
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Among other assumptions, consider

$$\sigma^2 = \begin{pmatrix} \sigma_1^2/2 & \sigma_\perp^2 & \cdots & \sigma_\perp^2 \\ \sigma_\perp^2 & \sigma_2^2/2 & \cdots & \sigma_\perp^2 \\ \vdots & \vdots & \ddots & \vdots \\ \sigma_\perp^2 & \sigma_\perp^2 & \cdots & \sigma_2^2/2 \end{pmatrix}$$





Conclusion

- Avoid relying on detailed model building, and take a coarse grained approach to the particle production in the early universe
- MEA captures the universal features arising from a Central Limit Threorem (concentration of measure)...
- ...as long as there's no hierarchy of couplings
- Break from weak scattering limit \Rightarrow Random Matrix Theory?
- Next: include expansion and metric perturbations

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Thank you